

\* Two Status

\* Syllabus Change! Covariance, Not. Univ. today.

Covariance in Special Relativity

- Why? • Need for Klein-Gordon & Dirac Equations!
- "Make your theories consistent with S.R.!"
- Preparation for study of General Relativity

Good References: Chewin "Classical Electrodynamics, 2. Ed."

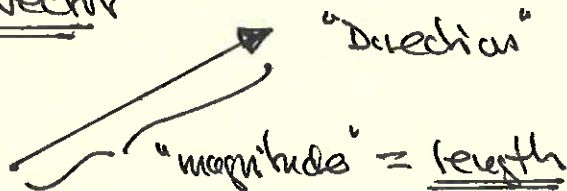
Vectors and Covectors

aka Contravariant Vectors and Covariant Vectors

aka "Shocks" and "Lasagnas"

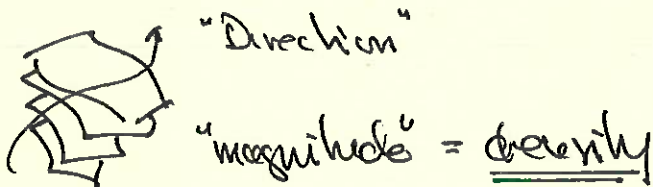
See posted ASP article: "Questions & Answers"

Shock Vector



Generic Notation:  $\vec{A}$

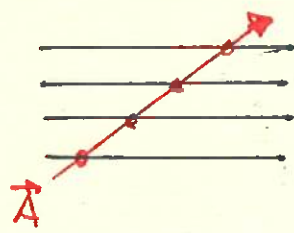
Lasagna Vector



Generic Notation:  $\vec{A}$

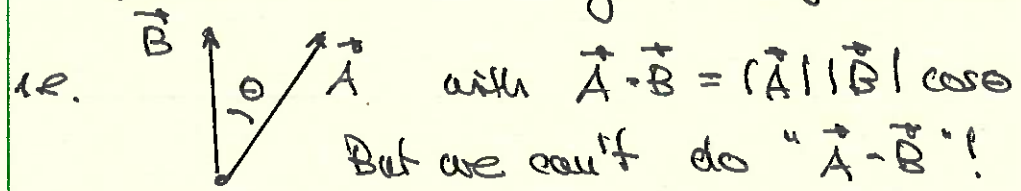
Prototype: Gradient of Equipotential Surfaces.

Inner Product: "Put the stick in the lasagna"



$\vec{A} \cdot \vec{B}$  = # of times the stick pierces a lasagna noodle  
 NOTE "UP" AND "DOWN" ARROWS!

So far we have no way to define an angle!



↳ Must invent some way to turn  $\vec{B}$  into  $\vec{B}$ !

Metric Tensor:  $\overset{++}{g}$  and  $\underset{--}{g}$  with  $\overset{++}{g} \cdot \underset{--}{g} = \underset{--}{1}$

i.e.  $\overset{++}{B} = \overset{++}{g} \cdot \underset{--}{B}$  and  $\underset{--}{A} = \underset{--}{g} \cdot \overset{++}{A}$

so  $\underset{--}{A} \cdot \overset{++}{B} = (\underset{--}{A} \cdot \underset{--}{g}) \cdot (\overset{++}{g} \cdot \underset{--}{B}) = \underset{--}{A} \cdot \underset{--}{B}$

Towards 4D Spacetime:  $\vec{A} \Rightarrow A^\mu = (A^0, \vec{A})$   $\mu=0,1,2,3$

"Coordinate Representation"

"Position" =  $x^\mu = (ct, \vec{r})$

Metric tensor  $\overset{++}{g} \Rightarrow g^{\mu\nu}$  w/  $A_\mu = g_{\mu\nu} A^\nu$  Implied Sum

General Relativity:  $g^{\mu\nu} = g^{\mu\nu}(x^\alpha)$

Special Relativity:  $g^{\mu\nu} = \eta^{\mu\nu} \equiv \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{bmatrix}$

"Minkowski Metric"

NOTE:  $\eta_{\mu\nu}$  is the same! i.e.  $\eta_{\mu\alpha} \eta^{\alpha\nu} = \delta_\mu^\nu$

## Special Relativity: Constancy of the speed of light

For all observers (i.e. observer #2 at  $V$  wrt #1)

$$d\tilde{t}^2 = c^2 dt^2 \text{ for an "expanding light front"}$$

$$\text{i.e. } ds^2 = \underbrace{c^2 dt^2}_{\#1} - \underbrace{d\tilde{t}^2}_{\#2} = c^2 dt'^2 - d\tilde{t}'^2$$

i.e. For position  $x^\mu$  need  $ds^2 = dx^\mu dx_\mu = dx^\mu \eta_{\mu\nu} dx^\nu$   
to be "invariant" =  $\eta_{\mu\nu}$  from next on.

↳ Define a four-vector  $A^\mu$  as an object that transforms so that  $A'^\mu A'_\mu = A^\mu A_\mu$  "invariant"

## Lorentz Transformation (for a vector)

$A'^\mu = \Lambda^\mu_{\nu} A^\nu$  keeps inner product invariant!

Not hard to show that for observer #2 moving at speed  $V$  along the  $x$ -axis, ...

$$\Lambda^\mu_{\nu} = \begin{bmatrix} 1 & -V/c & 0 & 0 \\ \frac{1}{\sqrt{1-V^2/c^2}} & \frac{-V/c}{\sqrt{1-V^2/c^2}} & 0 & 0 \\ -V/c & 1 & 0 & 0 \\ \frac{1}{\sqrt{1-V^2/c^2}} & \frac{V/c}{\sqrt{1-V^2/c^2}} & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}$$

You should check!  
Don't want to be  
class time for the  
exercise!

(for a covector)

$$A'_\mu = \eta_{\mu\alpha} A'^\alpha = \eta_{\mu\alpha} \Lambda^\alpha_{\beta} A^\beta = \eta_{\mu\alpha} \Lambda^\alpha_{\beta} \eta^{\beta\nu} A_\nu$$

$$= \Lambda^\nu_{\mu} A_\nu \Rightarrow \Lambda^\nu_{\mu} = \eta_{\mu\alpha} \Lambda^\alpha_{\beta} \eta^{\beta\nu} \quad (\text{inv})$$

### Important Four-Vectors

Position  $x^\mu = (ct, x^1, x^2, x^3) = (ct, x, y, z) = (ct, \vec{r})$

Velocity  $u^\mu = dx^\mu/d\tau$  where  $d\tau^2 = ds^2/c^2$

"Proper time" invariant!

$$= \left( c \frac{dt}{d\tau}, \frac{dx}{d\tau}, \frac{dy}{d\tau}, \frac{dz}{d\tau} \right)$$

$$= (\gamma c \beta, \gamma v_x, \gamma v_y, \gamma v_z) \quad \beta \equiv \frac{v}{c} \quad \gamma \equiv \frac{1}{\sqrt{1-v^2/c^2}} \quad v_i = \frac{dx^i}{dt}$$

Momentum  $p^\mu = m u^\mu$

NOTE •  $u^\mu u_\mu = c^2$  Prove it!!  $\Rightarrow p^\mu p_\mu = m^2 c^2$

$$\cdot ep^0 = mc^2 \gamma = \frac{mc^2}{\sqrt{1-v^2/c^2}} = E \text{ "Energy"}$$

$$\approx mc^2 + \frac{1}{2}mv^2 + \dots$$

Explicitly  $p^\mu = (E/c, \vec{p}) \quad \vec{p} = \gamma m \vec{v}$   
 $p_\mu = (E/c, -\vec{p}) = \eta_{\mu\nu} p^\nu$

NOTE: Quantum Mechanical Plane Wave  
 $\Psi(\vec{r}, t) = N e^{-iEt/\hbar} e^{i\vec{p} \cdot \vec{r}} = N e^{-i p_\mu x^\mu}$  Covariant!!

Why "covariance"?

- If an equation is "manifestly covariant" then it is trivially consistent with special rel.
- can build a theory based on all scalar, inner products

Calculus in 4D spacetime

Four-Gradient  $\frac{\partial}{\partial x^\mu} = \left( \frac{1}{c} \frac{\partial}{\partial t}, \frac{\partial}{\partial x}, \frac{\partial}{\partial y}, \frac{\partial}{\partial z} \right) = \left( \frac{1}{c} \frac{\partial}{\partial t}, \vec{\nabla} \right)$   
 $\equiv \partial_\mu$  covector!! HW!!

Also  $\partial_\mu \partial^\mu A(x^\nu) = 0$  is the wave equation!  
 $\Rightarrow$  ON TO • Electrodynamics (not us!)  
 • Relativistic QM (start Thursday)

Natural Units ( $\hbar = c = 1$ )

- What's special about "speed of light"? Nothing!!
  - we choose to measure  $t$  in "seconds", but we might just as well measure it in "meters"
  - think of  $c$  as a "conversion constant"
- If time: "Parable of the surveyors"

Little known fact: "c" has a defined value!

$c \equiv 299,792,458$  m/s (1983)

$\Rightarrow$  1 meter  $\equiv$  distance light travels in  $1/c$  sec  
 [we can measure "time" very precisely.]

So Drop the  $c$  altogether and set  $c \equiv 1$

[You can convert back later]

i.e.  $Mc^2 \rightarrow M$  in "energy units"  
 also momentum.

Typically we use MeV or GeV.



Example: Electron mass  $m_e = 9.1 \times 10^{-31} \text{ kg}$

$$\Leftrightarrow m_e c^2 = 8.19 \times 10^{-14} \text{ J} = 511 \times 10^3 \text{ eV}$$

$$= \underline{0.511 \text{ MeV}} \quad \text{Physicists know this, not this}$$

Do the same thing with  $\hbar = 1$ !

Recall  $\Delta p \Delta x \sim \hbar = 1 \Rightarrow$  Distance in  $\text{MeV}^{-1}$

$\Delta E \Delta t \sim \hbar = 1 \Rightarrow$  Time in  $\text{MeV}^{-1}$

$$\text{Useful: } \hbar c = 200 \text{ MeV} \cdot \mu\text{m} = 200 \times 10^6 \text{ eV} \times 10^{-6} \text{ m}$$

$$= 2 \times 10^{-7} \text{ eV} \cdot \text{m}$$

Example: Ground state Energy of Hydrogen

$$\text{From } 2 = \left[ \frac{2me}{-E} \right]^{1/2} \frac{e^2}{\hbar} \Rightarrow E = -\frac{1}{2} \frac{m e e^4}{\hbar^2}$$

We made this easier by using  $\alpha \equiv \frac{e^2}{\hbar c} \approx \frac{1}{137}$

$$\text{i.e. } E = -\frac{1}{2} m e c^2 \left( \frac{e^2}{\hbar c} \right)^2 = -\frac{1}{2} (0.511 \text{ MeV}) \left( \frac{1}{137} \right)^2 = -13.6 \text{ eV}$$

But we might have just said  $E = -\frac{1}{2} m e e^4 =$

$$= -\frac{1}{2} (0.511 \text{ MeV}) \left( \frac{1}{137} \right)^2$$

Thursday: Relativistic Quantum Mechanics!